

Celebrating 85 Years of Astronomy

Next Meeting

When:	Sat. Feb. 12th, 2022				
Time:	7:30 pm				
Where: See instructio	Online (Zoom)				
meeting on Page 8.					
	1 D				

Speaker: Dr. Joe Pesce

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Looking a little like the Starship Enterprise, the upper right formation is actually two gravitationally interacting galaxies. More details can be found at <u>www.sci-</u> <u>news.com/astronomy/hubble-trio-</u> <u>galaxies-10486.html</u>. Image Credit -NASA / ESA / Hubble / J. Dalcanton, Dark Energy Survey / DOE / FNAL, DECam / CTIO / NOIRLab / NSF / AURA / ESO / Judy Schmidt, www.geckzilla.com.

Star Dust

Newsletter of National Capital Astronomers, Inc. capitalastronomers.org

February 2022

Volume 80, Issue 6

Monsters in the Universe: New Insight into Black Holes

Joe Pesce NSF and George Mason University

Black holes are among the most enigmatic objects in the universe, and come in a variety of sizes, from several times the mass of the Sun to monsters with many billions of times the mass of the Sun (and maybe even atomic-sized ones). New advances in astrophysical instruments - with traditional electromagnetic observations as well as new ones using particles and gravitational waves - make this a golden age of black hole research. Recent findings from the electromagnetic spectrum, particles, gravitational waves, and the early universe are challenging traditional views.



This artist's concept illustrates a supermassive black hole with millions to billions of times the mass of our sun. Supermassive black holes are enormously dense objects buried at the hearts of galaxies. (Smaller black holes also exist throughout galaxies.) In this illustration, the supermassive black hole at the center is surrounded by matter in an accretion disk flowing toward the black hole. This disk forms as the dust and gas in the galaxy falls onto the hole, attracted by its gravity. Also shown is one of the two oppositely-directs jets of ultra-relativistic charged particles, believed to be powered by the black hole's spin. The regions near black holes are compact sources of high energy X-ray radiation thought, in some scenarios, to originate from the base of these jets. This high energy X-radiation lights up the disk, which reflects the X-rays, making the disk a source of X-rays. The reflected light enables astronomers to see how fast matter is swirling in the inner region of the disk, and ultimately to measure the black hole's spin rate.

There also exist stellar-mass black holes whose masses are about 5 to tens of solar masses, and intermediate-mass black holes, whose masses are hundreds or thousands of solar masses. Image Credit - NASA/JPL-Caltech (2013-02-21), <u>www.nasa.gov/mission pages/nustar/multimedia/pia16695.html</u>. Caption Authors - NASA/JPL-Caltech (2013-02-21). This file is in the public domain in the United States because it was created solely by NASA.

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Recent Astronomy Highlights

Small Supermassive Black Hole Discovered In A Dwarf Galaxy

Using the Chandra X-ray Telescope, astronomers have discovered a black hole about 200,000 times the mass of the Sun in a dwarf galaxy known as Mrk 462 which lies approximately 110 million light years away. The discovery holds the potential to help in understanding how black holes grow to be supermassive, either from stellar-mass black holes, or via direct collapse of a cloud of gas into such a black hole. If there are a large number of such supermassive black holes in dwarf galaxies, this would seem to favor stellar-seed growth process. The large ratio of high-energy X-rays vs. lowenergy X-rays seen by Chandra indicates that the black hole is obscured by dust and gas, a clue that perhaps many other such small supermassive black holes might lie in the hearts of other dwarf galaxies. More information on the discovery can be found at chandra.si.edu/press/22 releases/press 011022.html.

Observation of a Dying Star's End

In the Fall of 2020, a supergiant star in the galaxy NGC 5731, which lies approximately 120 million light years away, was observed to end its life in a Type 2 Supernova. Fortunately, the Pan-STARRS telescope in Hawaii detected an unusual amount of light coming from the doomed star in Summer 2020 allowing a team of astronomers running the Young Supernova Experiment (YSE) transient survey to watch the star, which was approximately ten times the mass of the Sun, over the period of its final 130 days. Such pre-supernova activity has never been observed and recorded before, with previous supergiants having been relatively quiescent up until the time that they went supernova. The observations will no doubt lead to advances in the understanding of the final stages of such supergiant stars lives. More information about the discovery can be found at phys.org/news/2022-01-astronomerswitness-dying-star-explosive.html continued on page 4 Abstract and Biography – continued from page 1



Biography: Joe Pesce is an astrophysicist whose primary area of interest is supermassive black holes. He has held research positions at the Space Telescope Science Institute and the Pennsylvania State University. He is currently a Program Officer at the US National Science Foundation, responsible for most of the US Government's ground-based radio astronomy facilities (the National Radio Astronomy Observatory - NRAO including the Atacama Large Millimeter/submillimeter Array (ALMA, in Chile), the Very Large Array (VLA, in New Mexico), and the US component of the Very Long Baseline Array (VLBA, spread across Earth). In addition to his day job, he is a Part-time Professor at George Mason University (Fairfax, Virginia) and a Visiting Professor at the University of Colorado (Boulder, Colorado). Previously, he founded, and served as CEO, of several firms specializing in high-level science and technology, critical thinking and problem-solving, consulting and education.

An important aspect of Joe's work involves public outreach, as an ambassador for science in general and astrophysics specifically. Among other activities, he has made numerous television and podcast appearances, and has served as science advisor to several science fiction series (television and online) and science fiction authors.

He is a Fellow of the Royal Astronomical Society and the Cambridge Philosophical Society; serves on the Board of Directors of The Presidents Leadership Class (University of Colorado, Boulder); is a proud alumnus of Peterhouse; is a member of the American Astronomical Society; the American Association for the Advancement of Science; Sigma Pi Sigma; the American Institute of Physics; and the Cosmos Club (Washington DC).

Joe received a B.A. degree in physics from the University of Colorado in Boulder, and M.Sc., M.Phil., and Ph.D. degrees in astrophysics from Cambridge University (Peterhouse) and the International School for Advanced Studies in Trieste, Italy.

Joe's other interests include science policy, space, leadership, interspecies communication, non-human intelligence, psychology, science fiction, and all things British. He conducts art-history research and collects art, antiquarian books, and antique furniture.

Exploring the Sky



"Exploring the Sky" is an informal program that, for over 70 years, has offered monthly opportunities for anyone in the Washington area to see the stars and planets through telescopes from a location within the District of Columbia. Presented by the National Park Service and National Capital Astronomers, sessions are held in Rock Creek Park once each month on a Saturday night from April through November, Beginners (including children) and experienced stargazers are all welcome—and it's free!

Hosted by: <u>National Capital</u> <u>Astronomers, Inc</u> and <u>Rock Creek Park</u>

Due to the ongoing Coronavirus Pandemic, Exploring the Sky sessions are canceled. When the situation changes, sessions will once again be scheduled.

More information can be found at NCA's web site, <u>www.capitalastronomers.org</u> or the Rock Creek Park web site, <u>www.nps.gov/rocr/planyourvisit/expsky</u> .htm. You can also call the Nature Center at (202) 895-6070. For general information on local astronomical events visit <u>www.astronomyindc.org</u>

The article-submission deadline for March's issue of Star Dust, is February 21st.

Clear Skies!

Fully Deployed James Webb Space Telescope Reaches Its Final Destination – A Halo Orbit of L2



Image Credit - NASA GSFC/CIL/Adriana Manrique Gutierrez

At the September 2021 NCA meeting, Dr. Heidi Hammel gave a fascinating talk about what the James Webb Space Telescope (JWST) would explore in the Solar System and the types of discoveries it was expected to make. (A recording of that talk is available at capitalastronomers.org/recordings/20210911 HHammel What-the-

JamesWebbSpaceTelescope-Will-Tell-Us-About-the-Solar-System.mp4.) But, no doubt on the minds of many who watched that talk were thoughts about all of the ways that the enormous space telescope might fail to deploy properly. Well, a little over four months after that talk, and a month after launch, at the end of what NASA dubbed the '29 Days of Terror', the almost completely deployed JWST fired its thrusters for about five minutes to change its velocity by about 1 ½ meters/second in order to enter what is known as a halo orbit around the Sun-Earth Langrange point known as L2. A number of other space probes, including the Wilkinson Microwave Anisotropy Probe, W-MAP, Planck and the Global Astrometric Interferometer for Astrophysics, Gaia, have also been sent to halo orbits of L2 for the advantages of being far away from the Earth and Moon, and their heat and fields, even while staying with Earth in orbiting the Sun, allowing for continuous communications with the spacecraft.

NCA member David Dunham has a long history with the concept of sending spacecraft to halo orbits of Lagrange points. He spent decades working with Robert Farquhar, known as the Father of Halo Orbits, who would jokingly describe their relationship: "Dave does all the work, and I take all the credit." In 1989 at a Goddard conference, Farquhar and Dunham gave a presentation first describing how a space observatory in a Sun-Earth L2 orbit had the advantage of access to at least half the sky, with the bright Sun, Earth, and Moon behind it. In 2021, David and four coauthors wrote a paper, "The Birth of Halo Orbits, a Fiftieth Anniversary", to memorialize the history of halo orbits and their use in space science and exploration. That paper can be found at iota.jhuapl.edu/AAS21-766paper.pdf. In addition, the pdf and audio file of a longer presentation David gave at a Future in-Space Operations (FISO) telecon are available at fiso.spiritastro.net/telecon/Dunham 11-3-21/.

February/March

Now in the morning sky, Mercury will reach Greatest Western Elongation in mid-February (see below). Venus also shines bright in the predawn sky. Mars will be in the morning sky. Jupiter is the only planet remaining in the night sky, low on the horizon at sunset until it begins to transition to the morning sky in early March, where it will join the other planets, including Saturn.

2/16	Mercury at Greatest Western Elongation – The planet will be 26.3 degrees from the Sun in the morning sky.
2/16	Full Moon at 12:38 EST

All times are in EST (Eastern Standard Time)

Occultation by the Trojan Asteroid (3548) Eurybates in the s.w. USA, 2021 Oct. 20 David W. and Joan B. Dunham

Eurybates is one of the Trojan asteroid targets of NASA's Lucy mission that launched only 4 days before this occultation. The Lucy mission is described in a good article, "Rock On", about it and other NASA missions to asteroids on pages 12-19 of the February 2022 issue of Sky and Telescope. On p. 17 is a box called "Get Involved", where the Southwest Research Institute (SwRI) asks interested amateurs to contact them, to possibly join their campaigns to observe occultations by the Lucy asteroids, and gives an example of one by (11351) Leucus that occurred in Arizona in late December 2019; Joan and I successfully ran four stations for that event, obtaining the southernmost chord and one other, as well as the constraining miss line on the south side.





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Editor: Todd Supple

Editorial Advisors:

- Michael Chesnes
- John D. Gaffey, Jr.
- Jeffrey Norman
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Thank you!

Recent Astronomy Highlights – continued from page 2

Strange Radio Transient Discovered

Using the Murchison Widefield Array (MWA) in Australia, radio astronomers have discovered an unusual source of radio waves. Sources that give off radio waves on a noncontinuous basis are called transients. Some examples of this phenomenon are objects such as pulsars, fast-spinning neutron stars from which radio telescopes receive signal in periods around that of milliseconds, and supernovae which can radiate brightly in the radio part of the spectrum for many days. But the source seen by MWA was extremely bright for about a minute out of twenty. It may be an ultra-long period magnetar, a slow spinning neutron star with a very strong magnetic field. More details can be found at

www.eurekalert.org/newsreleases/941140.

continued on page 7

Occultation Notes

- D following the time denotes a disappearance, while R indicates that the event is a reappearance.
- The times are for Greenbelt, MD, and will be good to within +/-1 min. for other locations in the Washington-Baltimore metropolitan areas unless the cusp angle (CA) is less than 30 deg., in which case, it might be as much as 5 minutes different for other locations across the region.
- Some stars in Flamsteed's catalog are in the wrong constellation, according to the official IAU constellation boundaries that were established well after Flamsteed's catalog was published. In these cases, Flamsteed's constellation is in parentheses and the actual constellation is given in the notes following a /.
- Mag is the star's magnitude.
- % is the percent of the Moon's visible disk that is sunlit, followed by a + indicating that the Moon is waxing and - showing that it is waning. So 0 is new moon, 50+ is first quarter, 100+ or - is full moon, and 50- is last quarter. The Moon is crescent if % is less than 50 and is gibbous if it is more than 50.
- Cusp Angle is described more fully at the main IOTA Web site.
- Sp. is the star's spectral type (color), O,B,blue; A,F,white; G,yellow; K,orange; M,N,S,C red.

Also in the notes, information about double stars is often given. "Close double" with no other information usually means nearly equal components with a separation less than 0.2". "mg2" or "m2" means the magnitude of the secondary component, followed by its separation in arc seconds ("), and sometimes its PA from the primary. If there is a 3rd component (for a triple star), it might be indicated with "mg3" or "m3". Double is sometime abbreviated "dbl". Often, rather than the separation, I give "dTime" or "dT", the time difference of the secondary star occultation relative to the primary star's occultation.

Sometimes the Axis angle (AA) is given. It is the angle measured around the Moon's disk, from the Moon's axis of rotation. It can be used with a lunar map to tell where a star will reappear relative to lunar features.

Mid-Atlantic Occultations

Vol 80, Iss 6

David Dunham

Asteroidal Occultations				
2022 2022	Day	EST/ EST	dur. Ap. Star Mag. Asteroid dmag s "Location	
Feb 10 Feb 11 Feb 14 Feb 15 Feb 17 Feb 18 Feb 19 Feb 21 Feb 21 Feb 22 Mar 7 Mar 12 *** [Mar 13	Thu Fri Mon Thu Fri Sat Mon Sun Sates Sun	3:12 4:56 22:36 0:22 5:39 2:33 22:54 23:17 3:26 5:59 0:01 1:02 and 22:59	4UC51652583 13.0 Eucharis 0.3 9 8 se-nVA,wMD;DC? 4UC34789304 13.4 Arenda 3.8 1.2 9 cOhio,nWV,nscVA TYC06300079 10.3 Walkure 5.2 1.3 5 CAZ,nAL,nGA,CSC 4UC55941483 13.7 Erato 0.7 15 10 MD,DC,nVA,nOhio 4U368130322 13.1 Radek 3.1 0.9 9 COH,nWV,nVA,SMD 4UC52347997 13.6 Gunlod 1.4 4 10 SNJ,C&nMDnVA? TYC07572115 9.7 Amalia 5.3 8 3 e-nVA,wMD,swPA 4UC51446027 13.6 Davewilliams3.5 2 9 SNJ,MD,nVA;DC? 4UC51940502 14.0 Bronislawa 1.8 12 11 MD,SNJ,SC-nWPA 4UC6150990 12.1 WilliamsBay 5.1 4 5 SNJ,nMD;DC,nVA? SA0 57301 8.8 Celuta 5.3 2.6 3 neTN,SWNC,SWSC 4UC52318483 11.9 Zhongolovich 5 2 6 COH,nVA,SDC,SMD times above are EST, those below are EDT *** 4UC51210550 14.4 Gyptis 0.4 9 11 NJ,PA,nOH;nMD?	
			Lunar Grazing Occultations	
2022	Day	EST	Star Mag % alt CA Location, Notes	
Feb 22 Mar 7	Tue Mon	7:38 18:53	alpha Lib 2.8 66- 22 -9N IAD,Vienna,VA;SDC;TH&Owngs,MD SAO 93315 7.7 27+ 48 8S YorkHavn,Lititz,S.Lansdale,PA	
		/	Lunar Total Occultations	
2022 Date	Day	EST/ EDT	Ph Star Mag % alt CA Sp. Notes	
Feb 12 Feb 13 Feb 13 Feb 17 Feb 18 Feb 20 Feb 20 Feb 20 Feb 22 Feb 24 Feb 24 Feb 24 Feb 26 Feb 27 Mar 7 Mar 7 Mar 8 Mar 9 Mar 10 Mar 11 Mar 13 Mar 13 Mar 13 Mar 13	Sat Ssan Frat Ssun Frat Ssun Frat Ssan Mon Tued Wedd Wedd Wedd Wedd Ssan Thui Ssan Ssan Mon Tued Wedd Wedd Ssan Thui Ssan Ssan Ssan Ssan Ssan Ssan Ssan Ssan	20:51 21:55 3:19 0:48 21:27 4:48 2:52 2:32 4:48 5:55 4:40 4:24 5:55 4:40 5:55 4:40 5:55 4:40 5:55 4:40 21:27 21:48 5:55 4:40 21:27 2:32 2: 2: 2: 2: 2: 2: 2: 2: 2: 2	D 39 Gem 6.2 88+ 74 77S F8 ZC1061, close double? D 40 Gem 6.4 88+ 76 19S B8 ZC1062, close double?? D ZC 1085 7.1 89+ 20 82N G8 mg2 7.8 sep 124" dT +4m R 46 Leonis 5.4 100- 65 56N M2 AA 269,ZC1544, TrmDst6" R ZC 1741* 7.1 93- 16 41S KO Axis Angle 214 degrees R ZC 1758 6.9 92- 47 67N G5 R ZC 1864 6.7 86- 38 45N K2 R 46 Vir 6.2 86- 47 60N K2 ZC1869,mg2 9,dTime+1.1s R 5 Librae 6.3 67- 36 30N K2 ZC105,mg2 10,dTim+2.7s R ZC 230* 6.7 56- 20 20N F3 maybe close double? R SAO 184545 8.3 44- 26 80N A7 Sun altitude -11 deg. R ZC 2544 8.1 33- 13 54S F0 Azimuth 129 degrees R ZC 2723 6.6 23- 2 47S KO Azimuth 129 degrees R SAO187318 7.7 22- 14 13N A3 Sun alt10, Az. 144 R omega Sgr 4.7 13- 16 59S G3 Sun alt. +3, ZC2910 D AG Arietis 8.2 19+ 14 59N B9 Az.275,ZC359,close db1 D SAO 93370 7.4 28+ 8 83S K5 Azimuth 286 degrees D SAO 93370 7.4 28+ 8 83S K5 Azimuth 286 degrees D SAO 76845* 8.2 47+ 39 81N A3 D SAO 76826* 7.9 46+ 40 39S A3 D SAO 76820* 7.5 46+ 44 60S F2 D SAO 76820* 7.5 46+ 44 60S F2 D SAO 76820* 7.5 46+ 44 60S F2 D SAO 76820* 7.5 46+ 44 03S F2 D SAO 76845* 8.2 47+ 31 52S B9 D 99 Tauri* 5.8 47+ 19 27S G8 ZC 742,mg2 12,dTime-20S D SAO 76885* 7.8 47+ 19 27S G8 ZC 742,mg2 12,dTime-20S D SAO 76885* 7.8 47+ 19 27S G8 ZC 742,mg2 12,dTime-20S D SAO 76885* 7.8 47+ 31 52S B9 D 99 Tauri* 5.8 47+ 19 27S G8 ZC 742,mg2 12,dTime-20S D SAO 76885* 7.8 47+ 31 52S B9 D 99 Tauri* 5.8 47+ 19 27S G8 ZC 742,mg2 12,dTime-4s D SAO 77533 7.3 56+ 54 84S G5 maybe close double? D SAO 776845* 7.8 47+ 40 39S A3 D SAO 77622 7.6 57+ 23 22N A3 times above are EST, those below are EDT *** D SAO 77610 7.2 75+ 19 11S F8 D SAO 80165 7.5 82+ 64 81N F2 ram so occultation light curves are sought.	
More information is at <u>http://iota.jhuapl.edu/exped.htm</u>				
David Dunnam, <u>dunnam@starpower.net</u>				

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Exploring the Sky

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Telescope Making

Guy Brandenburg <u>gfbrandenburg@yahoo.com</u> 202-262-4274 (leave message)

NCA Webmaster

Elizabeth Warner warnerem@astro.umd.edu 301-405-6555

Star Dust Editor

Todd Supple <u>NCAStardust@gmail.com</u> 301-595-2482 (h)

Social Media

Twitter: @NatCapAstro

Occultation by the Trojan Asteroid (3548) Eurybates in the s.w. USA, 2021 Oct. 20 – continued from page 4



Fig. 2. Observations of the occultation projected onto the sky plane at Eurybates. Credit: Marc Buie, SwRI (presented at the AGU meeting in New Orleans in December 2021). Red lines show the paths of the stations that recorded the star and had an occultation, with dots marking the disappearances and reappearances of the star. Blue lines show stations where the star was recorded but no occultation occurred, including our two lines, DD1 and DD2, underlined. Gray lines are stations where observations were attempted, but the star was not recorded due to clouds or equipment problems. The best-fit ellipse is in green, with axis lengths 58 km and 73 km. The path was accurately predicted because a few observations had been made during an earlier occultation. But it was not known then if Eurybates might have a large relatively close moon that could shift the paths of future events, including this one; that's why observers were spread out so far. The October observations showed there was no such moon, which will allow tighter coverage of future events by Eurybates, to better improve our knowledge of the asteroid's size and shape.

Another occultation, of a 10.4-mag. star by another Lucy Trojan, (15094) Polymele, will occur on March 27 UT (evening of the 26th EST) in the Carolinas, but will also be visible as far west as Kansas. The observation area will be decided only a few days beforehand when weather forecasts stabilize.

You are invited to join SwRI's and IOTA's campaign to observe that event. Information about it and other 2022 Trojan occultations is available at <u>occultations.org/publications/rasc/2022/nam22Trojanoccs.htm</u> while SwRI's Google Map for the March 27th path is at lucy.swri.edu/occ/20220327Polymele.html.

Recent Astronomy Highlights – continued from page 4

A Black Hole's Outflow Causing Star Formation in Host Galaxy

Black holes have a well-earned reputation for destroying objects that come too near to them, such as stars. But in one dwarf galaxy, a recent discovery indicates that black holes can actually help in the creation of stars. That galaxy, Henize 2-10, lies approximately 30 million light years away. There an outflow has been observed coming from a supermassive black hole that weighs approximately one million times the mass of the Sun. That outflow gas, traveling at around a million miles per hour, has collided with a cloud of gas 230 light years from the black hole, causing collapse of the gas to form stars. This is in contrast to other supermassive black holes, with plasma jets leaving at nearly the speed of light, so fast that they would destroy any clouds of gas in their path. More information on the discovery is at www.sciencedaily.com/releases/2022/0 1/220119194144.htm.

Calendar of Events

NCA Telescope Making, Maintenance, and Modification Workshop (TM3W) (previously the NCA Mirror- or Telescope-making Classes): <u>The</u> <u>Chevy Chase Community Center has reopened and classes have resumed</u>. Classes will be Tuesdays and Fridays, from 5:00 to 8:30 pm at the Chevy Chase Community Center (intersection of McKinley Street and Connecticut Avenue, N.W.) Please contact instructor Guy Brandenburg at 202-262-4274 (leave message) or at <u>gfbrandenburg@yahoo.com</u> if you plan to attend. Note that masks are mandatory, as in all DC government buildings. More info is at <u>guysmathastro.com</u>.

Open house talks and observing at the University of Maryland Observatory in College Park are temporarily suspended. When they resume, they will be on the 5th and 20th of every month at 8:00 pm (Nov.-Apr.) or 9:00 pm (May-Oct.). Updates are posted at <u>www.astro.umd.edu/openhouse</u>.

Next NCA Meeting: 12 March 7:30 p.m. Ira Thorpe (GSFC) The LISA Pathfinder Mission: Gwaves and Micrometeoroids

The APS Mid-Atlantic Senior Physicists Group: **(Zoom Meeting)** February 16th at 1:00 p.m., Jim Green, NASA, will give a talk entitled "The Exploration of the Moon and then onto Mars". More information on the meeting is available at <u>www.aps.org/units/maspg/meetings/meeting.cfm?name=SENIOR0222</u>. If you're interested in attending the meeting, please email <u>units@aps.org</u>.

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 Attending monthly scientific lectures on some aspect of astronomical observations Observing astronomical objects for personal pleasure at relat Attending large regional star parties Doing outreach events to educate the public, such as Exploring Building or modifying telescopes Participating in travel/expeditions to view eclipses or occultati Combating light pollution 	nomy ively dark sites ng the Sky ons			
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Are you interested in volunteering for: Telescope making, Exploring the Sky, Star Dust, NCA Officer, etc.?				
Please mail this form with check payable to National Capital Astronomers to: Henry Bofinger, NCA Treasurer; 727 Massachusetts Ave. NE, Washington, DC 20002-6007				



Celebrating 84 Years of Astronomy

Next NCA Meeting: 2022 February 12th 7:30 pm (On Zoom) Dr. Joe Pesce

To join the Zoom meeting, use the following link: umd.zoom.us/j/96856095178?pwd=cWhyNE92bGFYUkYxZ nl6eWVIK0IKdz09

Please download and import the following iCalendar (.ics) files to your calendar system: <u>umd.zoom.us/meeting/tJllcu-opz4rHdxfgBb8Lh5wRlgETFQ8lnI5/ics?icsToken=98tyKuC</u> upj4sGt2QsR6PRowAGo_4M_TxmCVcgqdFmhjHAXh_albh BO5FF4ZZIYDc

Please note that NCA Zoom meetings are often recorded.

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