Page 8

FROM HERE AND THERE -- Cont'd. from p. 3

We would like to print a list in STARDUST of the NCA members that belong to other astronomical groups such as the AAVSA, ALPO, the AMS and others. If you belong to one of these groups, please give the information to the secretary, Mrs. Nora Keel.

NEXT MONTH

9 9

0

8 ₹

Ŀ

ക

Our speaker for March will be Dr. David Musto, who will describe the beginnings of the U.S. Naval Observatory.

XXX

×ж

Astronomy compels the soul to look upwards And leads us from this world to another. - Plato

NO MD-DC JUNIOR MEETING in FEBRUARY.

Library, Naval Observatory Washington 25, D.C.

a manage and a state of the sta

FEB-1'68

$\varphi_{\mathcal{B}} \stackrel{\bigstar}{\xrightarrow{}} S_{\mathsf{T}} A_{\mathsf{R}} D_{\mathsf{U}} S_{\mathsf{T}}$

February's speaker is Dr. Joseph Weber, from the Department of Physics and Astronomy, University of Maryland. Dr. Weber's area of investigation recently has been gravity waves; he is in charge of experiments which promise possible detection of this elusive form of radiation.

Relativity predicts that an oscillating mass will radiate energy in the form of gravity waves, much as an alternating current will produce electromagnetic radiation. However, this radiation is on the order of only 10^{-97} watts for a terrestrial source^{*}, making detection virtually impossible. Even a spectroscopic binary star system does not radiate sufficient energy to be detected by means we now have at our disposal. Nature may provide a much more intense source of gravity waves; this might possibly occur when a star undergoes gravitational collapse. Conservation of angular momentum in the system would dictate faster and faster rotation as the radius shrinks, until the velocity of light is approached. Gravitational radiation energy of perhaps 10^{-9} watts would be attained. This could very well be detected with equipment now available to the physicist.

Dr. Weber's apparatus, which he will describe in his talk, has detected "events" which can apparently be explained only by brief pulses of gravity radiation, although it is still much too early to say defiinitely whether this is the case. Certainly it appears as though he and his staff are onto something big, possibly a development in physics fully as significant as the experiments of Hertz in the detection of electromagnetic waves emitted by an alternating gurrent. Fourt'd. p.2.

CALENDAR

FEBRUARY 3 6:00 p.m. Dinner with the speaker, Bassin's (14th & Pa.) Call Jerry Hudson, 948-2809 for reservations.

> 7:30 p.m. meeting of the officers and board of trustees. (Because of bad weather, we were unable to meet a quorum last month) Dept. of Commerce Auditorium.

8:15 p.m. GRAVITY WAVES, Dr. Joseph Weber, Univ. of Maryland. Dept. of Commerce Auditorium.

- 10 8:15 p.m. Discussion Group. (upstairs) Commerce Eldg., Rm.2062. Topic will be an amateur's observatory---that of our treasurer, Mr. Robert N. Bolster. Slides will be shown which were taken during various phases of construction, and mention will be made of certain Fairfax County zoning regulations pertinent to observatory building.
- 6,13,20,27 TELESCOPE MAKING CLASS at the Board of Education's Material Center in Bladensburg with Ted Noble.
- 2,9,16,23 TELESCOPE MAKING CLASS at the Chevy Chase Community Center at 7:30 P.M. with Hoy Walls.
 - 3 GENERAL MEETING OF JUNIOR DIVISION at 7:30 P.M. in the Dept. of Commerce auditorium. All juniors are urged to attend. January Junior Meeting was cancelled due to snow.

JUNIOR DIVISION ASTRONOMY CLASS Sunday afternoon at 3 P.M. in Planetarium of Montgomery Junior College. Call 933-0823 (Mark Goldberg) for the date.

* This was calculated by Einstein for a rotating rod 1-meter long, at an angular velocity as fast as the material would allow.

Page 2

GRAVITY WAVES - Cont'd. from p. 1

Dr. Weber received his Ph.D. in physics from the Catholic University of America in 1951. He has been the recipient of two Guggenheim fellowship under which he has done work in quantum electronics and general relative Before coming to the University of Maryland, he was with the Institute for Advanced Studies.

Those who may want to do some more reading on the subject will find an excellent, non-technical article in the October 1967 issue of <u>Scientific</u> Research (pp. 41-43).

.

MIDDLE EAST REGIONAL CONVENTION JUNE 14, 15, and 16, Pittsburgh, Pa.

The convention will be held at the Quality Courts Motel East (near the Westinghouse Bridge). It is on Route 30 off the Pa. Turnpike at the Irwin interchange. The motel is giving special rates. Friday evening will be a bus trip to the Allegheny Observatory. The Banquet will be Saturday evening with Dr. Wagman of the Allegheny Observatory being the speaker. Sunday will be a trip to large private Observatory in the Pittsburgh area. Plan now to attend this convention. In case you will consider presenting a paper, please write to Mr. George G. Lingbloom, 1606 Burchfield Road, Allison Park, Pa. 15101.

NEW MEMBERS

Membership Applications Received at the January Meeting

Joint Membership:

Darrel J. Freund and Darrel Jr. 4703 Teak Ct. Camp Springs, Md. 20031

wining a contra pro succe to through

NOTE FROM EDITOR

All of us associated with the publication of Stardust regret that it is often received late and sincerely hope that this has not caused you to miss a lecture that you particularly wanted to hear. The publication of Stardust requires the efforts of many people to collect the information necessary to write even just the calendar! It should be remembered that these people, including the editor, and the photographer are all voluntary and can not be expected to take time away from their regular jobs to work on Stardust. If you have never worked on a project like this you may not believe the number of things small and large that can go wrong--a last minute change in speaker, the speaker forgets to send his resume, someone is sick, someone is called out of town. When an emergency arises, there is no one prepared to take over and fill the gap. Several times in the last two years requests have been made in Stardust and at the regular meetings for help. No one has volunteered! Stardust belongs to the membership and your help is needed to make it the type of publication you want.

FROM HERE AND THERE

The Astronomical Society of Harrisburg will hold their annual dinner rebruary 15, Gallileo's birthday. Dr. Peter Van De Kamp will be the speaker.

How about the NCA having an annual dinner? Guess it would be our 31st year. Seems like it has been 10 or 12 years since we have had a banquet. Someone get the ball rolling.

We also note in the "STARDUST" of the Harrisburg group that their Junior group is selling the Messier OBSERVE Manual for \$1.50. Copies are still available from Bob Wright for \$1.00. - Cont'd.p. 3 LUNAR OCCULTATIONS CUMPUTED FOR WASHINGTON (LAT 38.920 LONG 77-065) FOR 1968 (DISTANCE FROM STANDARD STATION = -0 MILES) STANDARD STATION AT LAT 38.920 LONG 77.065

DATE	TINELUTI ACC H M S SEC	v 0	2.6.	USNO REF_NO	S.A.O. NUMBER	HAG	SP	РН	PERC	ENT	ÈĽG	PA DEG	CUSP ANGL	VA DEG	ALT	N AZ I	SUN AL T	AA Deg	LIBR. LONG	AT ION	I NIN MAG
CEB 10	6 7 9 20	в 2		Z0545B	77986	9.0	88	D	83	HAX	130	24	20N	322	39	277		.23	-5.3	-5.5	, _
FEB 10	8 32 50 15	83	i	205615	78108	8.5	F2	Ð	83	HAX	132	169	165	117	11	296		167	-5.4	-5.1	
_FEB 10	B 49 5 5.	.B_4		205648	76126	8-Z	KZ	.0	83	HAX	132	176		128	4	299		174	-5-4	-5.1	<u> </u>
FEB 10	73 72 41 8	6 1		Z06614	76648	8.5	AO	D	88	HAX	140	137	_ 54S	199	48	90	-9	132	-4.5	-6.3	
FEB 11	1 21 1 5	62	· ·	Z06699	78917	8.3	MO	0	89	HAX	141	95	84N	147	69	117		90	-4+8	-6.4	
FEB 11	3 26 52 15	71		206762	78965	B-2	AD	0	89	WAX	141	136	565	329	76	221 228		130	-5.2	-6.4	· ·
FEB 11	3 47 52 5	98	1066	207019	79141	5.6	A2	Ď	90	WAX	143	113	805	58	18	290		106	-5.9	-6.0	<u> </u>
FEB 11	9 43 37 28	9 2	!	207047	79164	7.8	_K0	^D	_ 90	WAX	144	184	95	134		298		177	-5.9	-6.0)
GRAZING	A 35 20 5	IEARU 9 6	1206	Z08057	79861	5.9	ко К	D	95	WAX	154	110	89N	73	72	224		96	-5.5	-6.4	
FEB 12	5 34 11 8	9.6	1211	208073	79869	6.Z	AO	D	95	HAX	154	163	385	109	63	248		151	-5.7	-6.3	i
FEB 12	9 23 32 16	91		Z08224	79980	8-0	65	0	96	MAX	155	40	18N	345	20 50	286		28	-6.0	-5-9)
_FEB 13	4 23 17 5	93	1334	209085	80529	7.0	Ğ5	Ď	98	WAX	166	130	845	138	73	171	• • • • •	Tiś	-5.1	-6.0	;
FEB 13	10 13 42 6	92	1357	209243	80634	7.7	60	D	99	HAX	166	154	625	100	18	282		138	-9.7	-5-4	• —
FE8 14	11 17 13 5	91	-	209990	98897	7.8	KÜ AR	R	100	HAN	170	229	15N 465	223	13	280	-9	207	-5.2	-4-5	j.
FEB 15	8 2 36 10	- 9 i		210604	99296	8.0	A3	R	99	WAN	166	250	59\$	214	53	zżī		229	-4.2	-3.6	; ·—
FEB 15	8 30 12 18	. 9 1		210625	. 99302	7.9	A0	0	99	HAN	166	202	115	162	49	235		181	-4.2	-3.8	<u>! </u>
FEB 15 FEB 15	84901B 847375	91	1576	Z10625 Z10630	99302	5.3	AU	8	99	WAN	166	324	425	281	46	240		303	-4-3	-3.1	2
FEB 17	6 35 13 7	93		ZL1926	136796	7.9	G5	R	69	WAN	141	270	685	290	46	153		248	-1.8	-1.3	
FEB 17	7 24 26 5	94	1790	Z11948	138813	7-6	MO	R	88	MAN	120	314	69N	321	48	170	-9	292	-1.9	-1.0	?
FEB 17	11 15 17 7	93	3 1901	212002	138861	8.0	85	Ř	68	WAN	139	337	. 46N	292	24.	244	5	315	-2.4	-0.4	
FEB 18	3 4 40 5	a 1		Z12542	139226	8.1	F2	R	61	HAN	129	301	81N	351	2	101		280	-0.5	0.1	i .
	3 38 11 7	73	1	212569	139243	9.0	65	R	81	MAN	129	345	33N 37N	38	10	106		329	-0.5	0.2	·
FEB 18	4 55 31 5.	់ខ្ញុំ	5	Z12593	139258	8.3	ĸś	Ř	81	WAN	128	324	58N	7	21	120		304	-0.5	ŏ.,	
FEB 18	4 49 43 7	73	1	212597	139261	9.1	60	R	81	WAN	128	346	36N	30	20	119		325	-0.5	0.3	1
<u>FE8_10</u>	<u>5 37 32 5</u>	<u>_9.</u> §	1911	712606	139268	4.0	F.U.	K	. 81	WAN	125	242	605	291	35	143.		242	Q+2. 		<u>, </u>
FEB 16	6 58 55 5	. 9 ŝ	í	212634	139282	8.4	кõ	я	80	HAN	127	297	85N	321	37	145		276	-0.7		í
FE8 18	7 9 23 5	<u> </u>	1	Z12641	139284	8.5	G5	R	80	WAN	127	323	59N	345	36	152		303	-0.7	0.1	
FEB 18	8 13 52 11	-9-3		732695	139305	9-3	60	R	79	WAN	126	285	825	260	36	212	• • • • •	264	-1.1		·
FEB 18_	11 28 5 9	ز ف_	j	2127.14	_157931	_8.7	_G5	Ŕ	. 79	HAN	125	247	. 44S	212	27_	227	-6	227	-1.3	_1.2	<u>. </u>
FE8 19	5 24 47 6	78		Z13404	158400	9.3	F5	R	71	HAN	114	339	418	23	13	121		321	0.6	2-1	
FE8 19	<u> </u>	7 6	2031	Z13414	158410	9.0	ĸö	R	70	WAN	114	276	765	310	18	128		258	0.5	2.2	2
FE8 19	6 27 56 5	7.4		213422	158418	9-6	60	R	70	HAN	114	287	875	323	. ZZ	133_		260	_0.5	_2.2	<u> </u>
FEB 19 FFB 19	0350 >	85 73	> 1	213424	158419	9+2	P8 45	R	70	HAN	114	282	825	290	23	135		264	0.5	2.3	
FEB 19	7 17 32 7	7 1	7	Z13442	158434	8.8	KO	R	70	WAN	114	266	665	294	28	144		248	0.4	2.4	,
<u>_F</u> EB 19	9 2 38 7	8 1	2040	213472	158453	8.6	KC	R	. 70	WAN		264	.645	271	35	172		246	_0.z	2,6	<u>ب</u>
FEB 19_	9.39.46	8.5	i	213486	156459	9.4	65	_ <u>R</u> .		HAN	113.	30,0	80N	_299_	_35_	182		282	-0-1	2.3	[
FEB 19	95706	86		713505	158472 158496	9-1	60 FA	R	69	WAN WAN	113	337	43N 48N	331	35	203		314	-0.1	2.6	
FEB 19	11 34 30 15	8 5		Z13540	158496	9.4		R	69	WAN	112	230	305	203	28	213	-5	212	-0.2	2.9	
	-2-31-442	7.4	2123	214266	183166	8.2	AQ.	R	-60	KAN.	101	301	. 76N	248	<u>_</u> 2	<u>це</u> .		286	يو با	ي ج	<u></u>
FE8 20	6 53 2 5	ล่ง	,	Z14312	183209	8.1	K2	R	59	HAN	100	325		- 4	14	131		310	1.5	3.0	3
FEB 20	8 45 38 5	θ 6	2168	Z14367	183253	9+1	F5	R	59	MAN	100	285	895	305	26	155	•	271	1.3	4.1	1
FEB 20	<u>8 42 10 6</u>	<u> </u>		214377	183257	9.2	<u>.</u>	- <u>R</u>	- 59	_HAN,	100	332	44N	323	26	154	-12	318	-1.0	4.4	
FE0_21_	7 24 50 5	7 6		Z15414	184150	8.5	FO	_R	48	MAN		301	_70N	342	_ī_	130		291	2.5	5.2	
FEB 21	7 40 23 6	8 6	1	Z15420	184161	8.5	F2	R	47	HAN	87	258	675	297	.9	133		24B	2.5	5.3	J.
FE8 21	8 27 14 5	86		Z15448	184185	9.0	A3			NAN.	87	283	<u>. ((N</u> 87N	315	15	141		273	2.4	بر يد 5.4	
FEB 21	8 50 6 5	7 7		Z15457	164195	8.8	F0	8	47	MAN	86	278	685	307	17	L46		268	2.4	5.4	<u> </u>
FEB 21 668 21	B 51 31 5	86	5	215461	184200	9.1	FO	R	47	WAN	86	303	67N 61N	332	17	146		310	2.4	5.4	2
FEB 21	9 56 24 10	7 4		Z15498	184230	9.6	K5	R	47	WAN	86	236	465	253	23	160		227	2.2	5.5	
FEB_22	8 53 25 6	7.6	<u> </u>	216571	185164	<u>9.0</u>	<u>_08</u> _	R	36	WAN.		314	49N	351	8	137		310	3.3	-6.4	
FEB 22	9 36 5 8	88 76	5	Z16540	185184	8+5	KQ 45	R	36	WAN	73	239	265 77N	305	19	158		282	3.1	6.6	
FEB 23	9 27 30 10	8 6	5	Z18339	186607	8.5	88	R	25	WAN	60	220	455	261	4	133		222	4.0	7.1	
FEB 23	10 19 34 5	9 6	<u> </u>	210367	186642	8.5	<u>A0</u>	_ <u>R</u>	25	HAN.	60	295	60N	329	-11	142		290	4.0	-7.2	
FE8 23	10 30 52 7	8.6	5	218404	186657	9.0	65	R	25	HAN	60	236	615	268	12	144		238	3.9	7.2	
FE8 23	10 42 2 7	8 !	5	218418	186666	9.2	X5	A	25	WAN	60	232	575	262	13	146		234	3.9	7.2	
FE8 23	10 48 52 5	8	<u> </u>	Z18420	186672	7.4	65	<u>R</u> 8	25	WAN WAN	60	289	ODN 6BN	315	14	148	-12	290	3.9		<u>'</u>
FEB 23	10 51 12 8	a	<u>.</u>	Z18432	186676	9.2	ĸo	<u>,</u> R	25	WAN	60	228	535	256	14	148	-12	230	3.9	7.3	
FEB 23	11 21 29 5	8 8	5	Z18451	186688	9.0	K2 89	A	25	WAN	60	282	738	306	17	154	-6	285	3.8	7.3	
FEB 24	10 5 31 6	0 3	í	220170	108013	8.6	65	R	16	WAN	47	233	665	276	2	129	-2	241	4.7	7.4	<u> </u>
FE8 24	11 11 7 8	. 8 :	5	220242	188066	9.4	MB	<u></u>	16	WAN	47	322	_24N	357	10	139	-8	330	4.6	7.5	<u> </u>
FE8 24	11 15 59 9 11 28 15 A	94	2831	220250	188073	5.0	83	R	16	HAN N	47	316	21N 30N	329	11	143	-7	325	4.5	7.5	1
FE8 25	11 34 42 6	9	1	221687	169239	9.0	K5	R	- 9	MAN	34	287	518	327	8	132		300	5.1	7.4	

۰.

page 6

OCCULTATION PREDICTIONS

These predictions are provided by David W. Dunham (Tale Observatory) and Thomas C. Van Flandern (U.S. Naval Observatory). The following explanations apply to column headings:

TIME (UT): Note that events are generally, but not always, in strict chronological order.

ACC SEC: Approximate accuracy of predicted time, in seconds. V: Value code, indicating expected usefulness of the observation. C: Observability code, indicating expected difficulty/ease. 9= easiest. Z.C.: Zodiacal Catalog number. A=double star (Aitken's catalog). B or C= other double star. D= double and variable star.

S.A.O. NUMBER: Star's number in Smithsonian Astrophysical Obs. catalog. PH: Phenomenon. G= graze. M'miss. It is possible for extremely shallow solid occultations to be listed as grazes, and vice-versa. ELG: Elongation, i.e., distance of Moon from Sun, in degrees. PA DEG: Position angle of the event, in degrees, meas. eastward from the north point of Moon's disc (not from the north cusp). CUSP ANGL: Distance of event meas. onto dark limb from N or S cusp. VA DEG: Vertex angle, meas. counterclockwise from uppermost (toward the zenith) point on Moon's disc (for use with altazimuth instruments). SUN ALT: Given only when critical. AA DEG: Aris angle, meas. eastward from Moon's north pole. MIN MAG: Magnitude at min. brightness of var. star (for which MAG=max.) GRAZING OCCULTATION NEAREN-toc.: This extra line is printed out if the star is within 0.02 radius (approx. 20 miles) of the mean limb at central occultation or closest approach. (This line refers to prev. line).

LUNAR OCCULTATIONS COMPUTED FOR WASHINGTON (LAT 33-920 LUNG 77-065) FOR 1968 IDISTANCE FROM STANDARD STATION = -0 MILES) STANDARD STATION AT LAT 38-920 LONG 77-065

			.								a ter	2.29.2					. عى	1.0210					<u></u>	<u>`</u>		
UA	ie –	· 11	71 E [(יניי	ALL SEC		U	2-6-		USNU	5-A-G.	MAG	SP	₽Н	PERC	ENT	ELG	PA	CUSP V	A	MOON	N 2	IUN AA	L18R/	TION NIN	ĩ.
				3	366					(EF NU	NUMBER				ZÓŶ	4,1 <u>7</u>		DEC	ANGL D	1EG (<u>ALT A</u>	121	L <u>T DE</u> G	LONG	LAT MAG	Ł
			E 2	20	-	-				+ - /				-					•							
34	1 21	- 22	26	20	6		2			14328	165304	9.3	K5_	<u>.</u>	6	WAX	27.		695_	<u>.59</u>	18	237	<u>_5_u</u>	4.4	<u>4.8</u>	
	0 21	23	53			5	2			224343	105310	7-9	A2	D	. 6	HAX	20	46	61N 3	358	8 2	248	67	4.4	4.7	
	8		20.		· •		2	2413		624983	140/90	- 9-1	12	Ų		WAX	39	93	_675	42_	_20_3	25.	U	0.0	_3.5	
Pt	0 1	23	23	4!		- 1	°			224985	146797	9.1	FB	D	11	WAX	39	72	885	27	20 2	245	94	4.0	3.5	
	8 Z		21	1	0	8	•			225002	146808	8.7	G5.,	D_	12	HAX	40,	. 25	. 45 🖵	135		260	43	<u>_3.9</u>	3.4	
FE	8 3	2	0	2	្ទ	_7	6			200256	109160	9.0	A2	H	19	нах	52	330	-8N 2	79	83	264	352	3.1	2-0	
	AZIN	GDC	ćúr.	TAT	ION N	(EAI	R B4	· /	AP:	PROXIM	ATE N.	LINI.	T LA	iT,∗	,= 3,8	. 69	- 0.	.06jW	. LONG	i. =	.71.0	261				
FE	B 3	22	50	33	5	- 6	4	1	A	200882	109616	8.6	GO	Ð	26	WAX	62	56	78N	26	51 2	219	-4 77	2.5	0+9	
F6	83	22	56	14	6	- 6	7			200866	109620	8.3	A5.	, D	. Z6	WAX	6Z	86	.725_	55	.50_3	221	-5 107		0.9	
F 8	8 3	23	35	35	25	- 6	2			Z00900	109630	8.7	A5	D	27	WAX	62	131	275	93	44 2	Z32	152	2.4	0.8	
F E	в э	23	52	15	25	- 6	7			200900	109630	8.7	A5	R	27	WAX	62	155	ر ډير	1,1,4.	42.3	237		2.3	0.0	
FE	8 4	Z	17	14	5	ē.	7			200962	109680	8.0	K5	D	27	HAX	63	43	65N 3	352	16 2	265	64	2.0	0.6	
FE	8 5	1	46	. 4	6	- 6	5			Z01611	92849	9.Z	KD	D	36	HAX	74	28	. 49N_3	37	33 3	25,8	47	0.9	-0.7	
FE	8 >	2	15	43		- 6	7			201623	92656	8+7	K5	0	36	KAX	74	60.	81N	B	28 2	263	79	0.8	-0.8	
. FR	B 5	. 3	58	_ 6	10		3			Z01658	92671	a. z	AU	9.		NAX	75	. 9	30N 3	suz_	83	279_		0.6	-0.8	
FE	8 7		12	31	~ ~	- 2	2	486		202748	75886	5.2	83	U	55	WAX	95	50	65N	3	60 2	241	64	-1.6	-3.3	
FE	8 7	2	59	1	21	- 1	9		A.	202764	75906	7.4	65	D	55	WAX	96	157	.95_1	01	40 3	265	171		-3.4	
FE	8 7	3	46	46	6	6	?			202787	75923	0.0	HI	D	56	WAX	96	106	60S	49	32 3	272 .	- 119	-2.0	-3-4	
. Ft			37	38	2	•	2			ZUZBUJ	75930	6.9	- FB	D.	56	WAX	.97	70	BAN	14	Z 3	280		-2.1	-3.4	
FE	8 8	1	20	22	, e	<u></u> ?	1			203476	70491	8-8	52	0	. 64	WAX	105	103	675	58	68 2	234	- 112	-2.6	-4.4	
F			35		2	e.	6	••		203478	76493	6.3	- KQ.	<u>P</u>	- 04	WAX	106	74	84N_	27	67	238		2.6	-9.4	
FE	в в	2	21	28	- 2	- 2	2			203499	76509	9.0	60	5	65	HAX	107	54	63N 3	157	53 2	259	64	-3.1	-4.5	
	6 8	2	21	27		- 1	- 1			201042	16537	8. /	88	.0.	65	MAX	108	131	405	74_	. 25 . 3	282	140	-3.4	-4.5	
FE	8 8	2	26	11		9	1			203242	76538	8.0	KU	D	65	HAX	108	48	57N 3	352	Z4 3	583	57	-3.5	-4-4	
FE	8 8		53	10	15	8	8	63Z		203549	76542	8.2	AZ	0	66	HAX	108	14	Z3N_3	19	19	287	2;	-3.5	-4.4	
FE	5 8	22	59	33	8	- 1	2		Α.	204060	76895	7.8	85	0	73	WAX	117	119	575 1	75	63 1	107	-5 ,12:	-3-2	-5-2	
FE	8 8	_ 23	ιo	28	- 2	- 9	9	750/	Α.	204067	76903	6.9	_F5_	<u> </u>	73	WAX	<u>, U</u> 7	_9Z_	B451	47	65	ш_	-7 97	-3.2	-5.2	
F	8 8	23	12	13	<u> </u>	- 1	. 10			204068	16404	. O. B	65	0	73	WAX	117	97	795	152	65	112	-7 102	-3.2	-5-2	
F	8 9	1	. 29	40	2		1			204102	76936	9.3	KU	<u>P</u>	73	HAX	118	. 21	25N_3	53_	- 10 - 1	212	X	2		
F	B 9	<u> </u>	31	2	8	- 4	9			204110	76945	7.1	A2	D	73	MAX	110	132	445 1	104		212	13	-3-8	-2-3	
- FE	B 9	1	19	- 5	2	- 4	8			Z04111	76946	8.9	۲Z	O.	73	WAX	110	.88	885 -	<u>. 10</u>		201		<u> </u>		
FE	8 9	2	51	32	5	5	8			204137	76965	8.3		0.	74	HAX	118	80	83N	24	04	251	8	-4-0	-5.4	
FE	8.9	. 5	8	46	23	- 6	2			204177	76999	. 9. 0	_ K5_	. <u>p</u> .	- 74	WAX	115	168	951	08_	18	276	17	-4.4		
FE	8 9	5	1	19	- 11	- 6	7			204178	77000	9.0	65	0	74	HAX	114	124	235	93	39	215	158		-2.3	
	0 9	- 7	41	45	8	B	6			204253	77061	9.0	GU	U	15	HAX	120	40	42N	148	-11-1	296	43	<u></u>		
FE	89	7	51	25		8	7			204265	77070	8.0	65		75	WAX	120	45	47N 3	122		298	45	-4.0	-5.Z	
FE	B 10	0	9	50	. 10	4	. 4			Z05180	11111	8./	68	D	81	XAW	129	20	170	76	- <u>Þ.(</u>	110		Z	- 0 · D	
FL	B 10	. o	9	- 7	10	- 5	з			Z05187	77783	9.1	AZ	0	81	HAX	129	139	445	194	01	111	139	-4-2	-6-0	
FI	8,10	0	27	35	. 0	_5	3		· · · ·	205207	77801	9.0	60	. <u>P</u>	81	MAX	129	134	495	86.	10	117	134	-4.Z	-6.0	
FI	EB 10	1	. 3	31	6	- 4	. 0			Z05240	77818	7+0	×>	D	81	WAX	129	- 57	24N	97	16	135	57	-4-4	~6.0	
F!	E 10	Z	14	34	e_	- 5	2			205288	77853	9.0	40	D	82	WAX.	129	40	36N	16_	18	200		- 4.0		
FI	EB 10	2	25	31	6	- 6	2			ZD5303	11866	9.0	KO	D	95	MAX	129	60	50N	29	- 11 - 3	216	60	-4.7	-0-0	
FI	B 10	2	51	29	6	. 6	2			Z05315	77875	9.0	К2	Ð	82	нах	129	41	37.0 2	55	. 74	234			_0.0	
FI	B 10	4	46	33	6	6	1			205404	77944	9.4	KO	D	82	HAX	130	125	59\$	64	23 2	266	124	-5.1	-6.0	
F1	B_10	4	_52	29	5	6	3			205418	77957	8.6	. AQ	9_	82	HAX	130		905	32	32 2	267		-5-1	-0-0	
F	EB 10	- 5	18	49	9	6	2			205420	77959	9.3	FØ	0	82	WAX	130	155	295	93	47 2	271	154	-5-2	-6.0	
FI	EB 10) 5	5 39	34		8	4			205457	77967	8.5	B9	D٠	82	на х	130	_100_	845	39	43 3	Z.74	99	-2+2	~ 5.9	

JANUARY LECTURE - ROCKETS, ASTRONOMY, AND AERONOMY

X-ray stars are the hottest topic in astronomy today according to our January speaker--Mr. Charles Johnson-- head of the Aeronomy Section of the Naval Research Laboratory. The most famous X-ray source is the Crab Nebula (Taurus XR-1) which an Aerobee rocket flight proved several years ago is not a point source of X-rays. This flight occurred during an occultation of the nebula by the moon, and the X-ray emission did not cut off sharply but decreased gradually. The strongest X-ray source is Scorpius XR-1. Most X-ray sources exist along the galactic equator with the greatest concentration toward Sagittarius. Few are extra-galactic; a notable example of one of these is quasar 3C-273. One variable X-ray star has been discovered; i.e., Cygnus XR-1, whose strength decreased significantly between June 1964 and the following April.

Aeronomy is a new field of science barely a decade old which deals with that part of the upper atmospheres of planets where ionization and disassociation are important. Here the density of air is millions of times less than at the surface.

The German V2 rockets were the first to observe and record the ultraviolet spectrum of the sun. Only by rocket or satellite can the sun's Lyman alpha line of hydrogen (1216 A.) be seen. After the V2's had been used up, N. R. L. Aerobee rockets continued upper atnosphere research by carrying a wide variety of instruments into near space up to 250 kilometers. Various bands of ultraviolet can be selected and measured by detectors having different combinations of windows and gases; for example, nitric oxide exposed to uv through a lithium fluoride window. The window sets a lower limit on the detected radiation wavelength and the gas, an upper limit. A Schmidt camera with a calcium fluoride corrector plate and image intensifier electronics has been used on board Aerobees for studying hot 0 and B type stars in ultraviolet light down to 1250 A. These optics are blind to Lyman alpha radiation which is strong in air glow and would fog uv star fields.

The Bennett R F mass spectrometer has been successfully used on Aerobees for measuring ion concentrations in our upper atmosphere. This instrument has revealed many interesting facts about the ionosphere such as the presence of a great deal of the nitrous oxide ions in our upper atmosphere. (Perhaps there is a Martian Dr. Kiess who thinks that life is impossible on the earth because of the presence of this poisonous compound of nitrogen in our atmosphere!) The Marina V fly-by of Venus studied the ionosphere of this planet. Apparently the ionosphere of Venus is high at night but low during daytime. Mr. Johnson explained that since Venus has no appreciable magnetic field, the solar wind is strong enough to blow the ionosphere off the planet or to the night side.

- Leith Holloway

FROM HERE AND THERE ... Cont'd. from p.2

From "THE GUIDE STAR" of the AAA of Pittsburg: How are we going to teach logic in a world where everybody talks about the Sun setting, when it is really the horizon rising?

"STAR LITE" is the name of the newsletter sent out by the Peoria Astronomical Society of Peoria, Ill.

"STAR Newsletter" is the title of the publication of the Society of Telescopy, Astronomy and Radio (STAR) of Eatontown, N.J., a society belonging to the Middle East Region of the Astronomical League.-Soncluded on p. 8

The January-February issue of Review of Popular Astronomy carries an article by Jerry Hudson. It is the same article that appeared in a recent STAR DUST on A Simple Clock Drive Design.





Before one has gone once around the barrel, the HCF has worn into contact with the optical surface.

HCF is useful too when making gross zonal corrections, such as removing a severe case of turned-down edge, or correcting a hyperbola. Final smoothing and blending can only be done on a pitch lap, however.

This suggests one might wish to use another piece of glass (preferably one with approximately the right curvature) for the pitch lap, keeping the original tool free for HOF working. One can proceed a bit more confidently on his pitch lap, knowing he has an intermediate alternative to the one dreaded by all mirror-make returning to find grinding.

OCCULTATION PREDICTIONS

the use of the table is self-explanatory, with pessibly the following explanations needed:

Z.C.= Zodiacal Catalogue number. If followed by "N", take the mean time of the appearance (disappearance) of the double star. For any other letter, use the brighter component of the pair.

Some Z.C. numbers have been assigned to bright Solar System objects:

Mercury	4001	Uranus	1007
Venus	4002	Neptune	1008
Mars	4004	Ceres	5001
Jupiter	4005	Palles	5002
Saturn	4006	Juno	5003

PHEN- Phenomenon D- Disappearance R-Reappearance *=Near Graze P.A. DEN-Position Angle (degrees), measured eastward from the northernmost point on the moon's disc. For example, 90° would be the center of the dark limb during waring; 270° during waring.

P.4. C.B. La-Position Angle of the Center of the Bright Limb. (same convention) Altitudes of the sun and moon are only given when these are important.

LUMAR DICLUTATIONS COMPUTED FOR MASHIMETON - [127: 38.920.110% - 77.0531 EOR 15982 (DISTANCE FROM STANDARD STATION - _ O MILES) - STAMDARD STATION AT LAT 98.920 LONG 77.065

DATE	. Т.П Н	HECUI M	s se	<u>_</u> x	0.	2	USNO REF NO	S.A.D.	_ MAG	SP_6	<u>NH PER</u> SL	ICENT	ELG	PA I	LUSP ANGL	VA DEG	MOON ALT AZ	SUN	AA Deg	LIBRA LONG	LAT MA
JAN	22	42 5	0	57	9	3143A	223031	190356	7.5	KŌ	0 10	HAX	36	22	3 3 N	347	18 22	3 -9	39	5.6	6.6
L <u>IAN 3</u> JAN 3	22	36.5	i31. !3	0. <u>5</u>	. 6		ZZ3912	165002	9.0	<u>K2</u>	D 17	LHAX.	48. 48	.113_ 75	505. 885	<u>86</u> . 45	27 21	9 -11	<u>133</u> 95	_ <u>عدد_</u> 5.7	5.0B
	23	47 4	ņ	5 5	<u>,</u>		723929	165016	8.6	65	يت_ق	WAX	48	37	54N	359	20 22	2	57	5.6	5.7
JAN JAN 4		19.1	4	56	<u>.</u> .		223935	165024	9.0	. 65	<u>د م</u>	RAX	-49	_62	7.9N	<u>_ 21</u>	15.23	5	دف_	5.6	
JAN 4 GRAZIN	ן ובםם	11 4 Сш.та	6 3 T.10N	9 6 NFA	9 RHY	3284	223961 PROXIM	16503Z	7.1 LIHU	F5 [LA1	D 11	WAX	- 0.	140 .35tw	235 . LOI	93 16	7 24 77.06 -	4 }	160	5.5	5.5
JAN 4	1	19 5	6 3	8 6	9	3764	223961	165032	7.1	F5	R 17	WAX	49	155	75	108	5 24	6	175	5.5	5.5
JAN 4	23	11	.3	5 5	7		Z24609	146517	6.6	60	D 53	WAX	60	41	6ZN	16	36 21	1 -12	6Z	5.5	4.7
AN AL	22	22 4	8 7	55.	- <u>8</u> 7		724610 724631	146516	4.8.8	. <u>FQ</u> G0	02 D 25	S WAX	. <u>60</u> . 60	<u>60</u>	53N	354	26 Z3	<u>9 -10</u> 1	<u>07</u> 54	5.5. 5.3	4.6
JAN	<u>ï</u>	-21-1	6	ē	<u>.</u>		124649	146547	9.1	K5	0.20	RAX	لغ:	104	555	59	17 24	3	125	5.2	4.5
JAN 5		34_	7	5_6	1		125239	146972	0.8	FB	<u> </u>	L.HAX	.72	62	_85N	. 27	35 22	8	85	4.6	<u></u>
JAN 6	2	22	2	57 87	7		725273	146992	8.8.	F8 F0	D 35	S WAX	73	55 120	78N 47S	8 59	18 25	1	132	4.3	3.2
JAN	22	29	2 1	25	6		200462	109303	9.0	FO	D 44	WAX .	83	112	455	116	52 17	6 -6	134	4.1	2.3
JAN	0 1	33 4	5	5 5	6	62	200483	109319	9.1	F8	D 44	HAX	83	43	66N	12	45 22	2	65	3.6	2.0
AAL	<u></u>	3 5	53	<u>6 5</u>	6	A	200489	109325	9.9	<u>_F8_</u>	D 44	HAX	<u>83</u> 83	17	40N 80N	<u>339</u> 21	41 23	0 A	<u>38</u> 78	3.5	2.0
JAN	š		ý	ś6.	<u> </u>		100565	109385	. O. A	ĸź.	P4	. HAX		. 63	745	<u> </u>	16 26	i	104	3.2	
JAN JAN		34 25.2	33 10	7.5 1].	ŝ.		200587	109399	.8.8	6 F5	0.4	I NAX	. 85	105	525 42N	328.	6 26	9	41	3.1	1.7
JAN .	22	28 1	0	5 3	97		Z01139	109802	7.8	65 65	D 53	NAN C	94	70 84	885	88	56 15 58 17	7 -5	91	2.9 2.8	0.9
J AN 8	i î	10 2	3	5 5	7		201163	109827	8.9	65	D 54	WAX	94	68	905	34	51 22	5	68	2.3	0.6
. JAN E Jan e	1.2	40.1	.6 10	8.5. 76	. <i>9</i> . 7		201217	109853	8.6	K2	0 51	E.WAX	95	110	495. 405	<u>. 62</u> 58	13 27	0	131	1.8	0.3
JAN8		18	2	5B. 5 6	- 6.		201286	.109906	.8.9	.K0	05	L HAX		94	645. 66N	103	7.27	5	.115 84		-0-4
JAN	1	.13	3 1	<u>- 6</u>	ì.	305	201836	_92788	B.7	_KÇ	0_6	HAX	105	124	365		_60_21	<u>6 </u>	143		-0.8
JAN SAN SAN SAN SAN SAN SAN SAN SAN SAN S	, 1 . 1	.513	15	5 6 56	. <u>9</u>		201838	92789	8.1	Р5 К0	D 63	NAX A	105	64	84 N 		55 23	í	82	0.9	-0.8
JAN 4	4	35 2	23	67	7		201689	92814	6.7		0 6	MAX	106	41	61N	348	26 26	7	59	0+4	-1+0
JAN	6	11	4 1	2 9	2	325	201925	92837	7.9	KO	0 6	WAX	107	136	255	84	8 20	1	154	0.3	-1.1
JAN JAN	22	25 4	13	5 7	5	A	202327	92840 93112	9.2	_KC	0 73	Z WAX	115	43	60N	90	54 11	7 -5	59	0.3	-1.6
JAN3	22 22	-31_1	9	53	- 9-		202331	93115	7.4	M6	D7	A HAX	.115		86\$	124	-55-11	<u>aa</u>	- 93	4.0	<u>-1.8</u>
JAN 14	s	. 922	25i	59		<u>A32</u>	202419	93189	5,9	MB	0 .7:	S. WAX	117,	141	235	86		Ś	156	-0.9	-2.5
JAN 10) 6)7	40 2	25	59 59	6		202466	93225	9.2		D 74	L HAX	118	95 166	695	42	14 26	3 1	110	-1.1	-2.4
GRA7 U	e Uč	<u>. UL 7</u>	LON	NEA	RRY	N	202021	ATE S.	LÍĤÌ	T 143		19.04	+		I Q	NG	77.06	1	-;;		
JAN 10	<u>23</u>	_#3_3	1 9 35	a	1		702944				.D	1.HAX	127	.111_	575	150	62.12	L	124		-3.2
JAN 1 JAN 1		48 3	85 1 23 1	27	6		202960	76050) 7.8 6.8	A0 A2	D 6) MAX Lihax	127	126	42S 14N	150	71 15	2 8	136	-1.3	-3.3
JAN 1	1	50		5 7	4		202975	76069	6.7	K2	0 80	XAN (127	42	53N	27	73 19	7	54	-1.5	-3.4
JAN 1		26	232	5 9	5	540	203070	76154	6.0	AC	D B	L WAX	128	- <u>1</u> 56	67N	356	38 27	0	68	-2.2	-3.6
GRAZ L	NG 00	158. CULT	58 A T 1 CM	59 NEA	R81	- A	203111 PPROXIM	ATE N.	LINI	T LA	<u>G_6</u>	2.HAX 38.93	129	- 351 • 3 1 (A	2N	295 NG	- 77.06	3	2	-2.4	
JANL	<u> </u>	10	46	5 9	4		703144	76224	64	<u>89</u>	DB	A HAX	129	64	75N	9	19 28	4	76	-2.4	-3.6
JAN I	i;	40	۰ • ۵ ۱	õ6	. A.		203163		6.9	.89	ه_ه	L.RAX	129	.141_	295.		13.28	a	152	-2.5	-3.6
JAN I	1 7	46	39 52	79	15		203175 203178	76254	B.C	F5 F5	08 D8	2 WAX 2. WAX	129	125	45S 66S	72	12 28	9 0	136	-2.5	-3.6
JAN 1	1 6	40	30	6 7	4		203200	76283	5 7.8	60 40	D 8	AX WAX	130	119	515	71	3 29	-10	131	-2.5	-3.5
JAN 1	2 1	27	6 1	9 7	r i		203704	7663	7 6.7	G5	DB	7 WAX	138	143	325	176	73 14	0	151	-2.4	-4.5
	2] 2	. <u>59</u> . i 31	50 34	78 6 9	1.2	698	. <u>703719</u> 203761	76651	5_B_2	KZ	08	B_RAX B NAX	139	-116-	<u>595</u>	351	<u>76 16</u> 46 26	5 6	124 56	لکیة:	-4.5
JAN_1	<u>.</u>	5.43.	55	<u>9</u> 9	i.		204.52	77251	3.8.0	_KO	0.9	A WAX	151	151	345	69	45 27	h	153	-4.3	-5.4
JAN L Jan J	4	27	20 59	6 9	52	-	20584	7829	L7.5	K2	0 9	7 NAX	161	103	825 39N	115	57 9	8	51	-3.8	-6.0



POLARIS TELESCOPES

Many portable telescopes lack a ready means for lining up the polar axis with the earth's axis, thereby assuring the observer of hours of tracking by the clock drive, and making possible the taking of long exposures. Either the problem is ignored, whence observing becomes less pleasant and photographs unsatisfactory; or the observer spends half an hour tediously lining up the mounting, using his finder for sighting.

A glance of Fig. c shows why even careful guiding is not able to eliminate trailing of star images if the axis is misaligned. The film and camera remain in the same alighment north-south with respect to the false pole, F, while the direction on the film toward true north, P, changes during the exposure. Of course, the guide star remains fixed on the film, while all other parts of the field rotate about that point. The angle through which the film rotates is $\Theta_1 = \Theta_2$. Under the worst possible conditions, for a pole misalignment F, a co-declination of the guide star D (90° - decl.), and an exposure of length T.

$$(\theta_1 - \theta_2)_{\max} \approx \frac{EWT}{D}$$

W is the angular rate of the earth's rotation, 15 /hr. Thus, if our polar axis is misaligned 2°, our guide star 30° from the pole, and we make an hour's exposure, we could rotate the field as much as 1°, which would easily be seen on any photograph. Cameras of short focal length are not immune to this effect-all that matter are film size and off-set of the guide telescope.

The polaris telescope shown in Fig. a is patterned after one built by Mr. Robert N. Bolster, and could be adapted to any standard mounting. A means for adjusting the alignment of the telescope is needed, such as the ring mounts shown. The reticle is a standard cross-line ruled on glass (obtainable from Edmund Scientific Co.). A ring of the correct angular radius (54' for 1968.0, +0'. 3 per ann.) is scribed onto stiff, transparent plastic with a pair of dividers. One should use very light pressure, and practice on several scraps before doing the final one. The plastic is then centered on the cross-line reticle, and the two mounted inside an eyepiece. The threaded retaining ring that was intended for the field lens can be used; machine washers can be inserted to obtain the right focus.

In use, the mounting is lined up approximately, and the main telescope pointed to any of the following bright stars:

- I. (0^h) & Androuedae (N.E. corner of the Great Square)
- II. (6^h) & Orionis (Betelgeuse)
- III. (12^h) Y Ursae Majoris (in the Big Dipper)
- IV. (18^h) Y Draconis (in the Dragon's head)

Final adjustment is made by lining up the Polaris telescope with Polaris on the appropriate hair and the scribed circle, as shown in Fig. b. The view is as one would look directly into the eyepiece in Fig. a, and it assumes a simple star diagonal was used in the Polaris telescope.

Adjusting the rig consists of bringing the intersection of the cross hairs into coincidence with the rotation of the mounting and turning the reticle until hair I coincides with an angle $30^{\circ}(2^{n})$ east of the meridian. The main telescope should be clamped to point along the meridian.

A little time in the shop making a Polaris telescope will be repaid during the first few nights of observing. <u>Lining</u> up should not take over five minutes, and accuracy can be held consistently to within 1/10 degree.

